

What stellar populations can tell us about the evolution of the mass–metallicity relation in SDSS galaxies

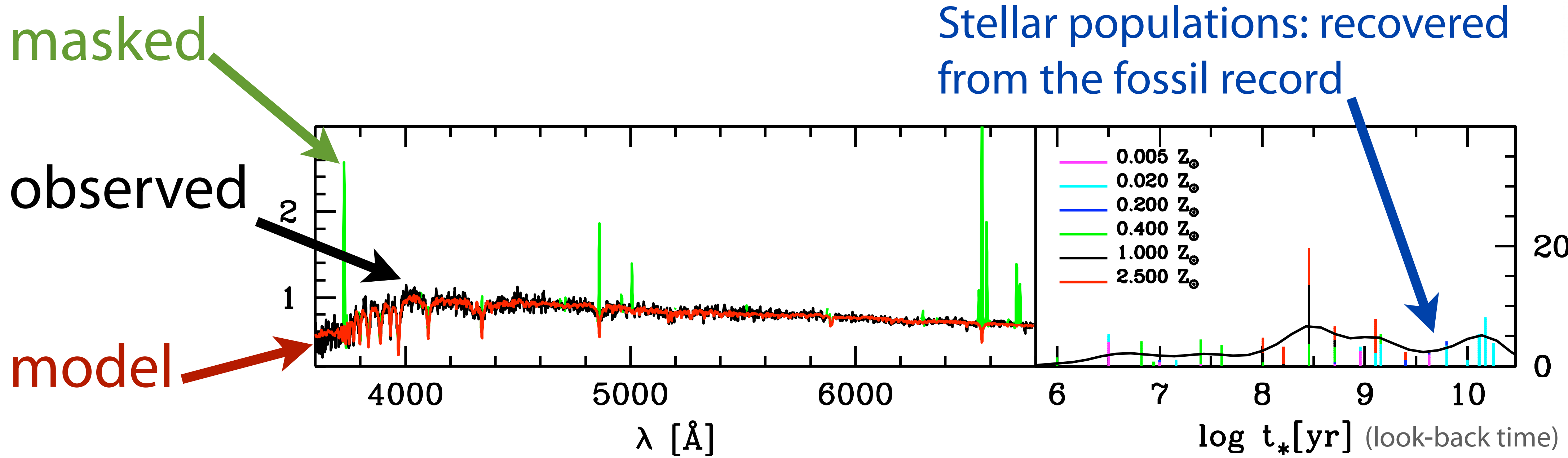


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Astropaleontology

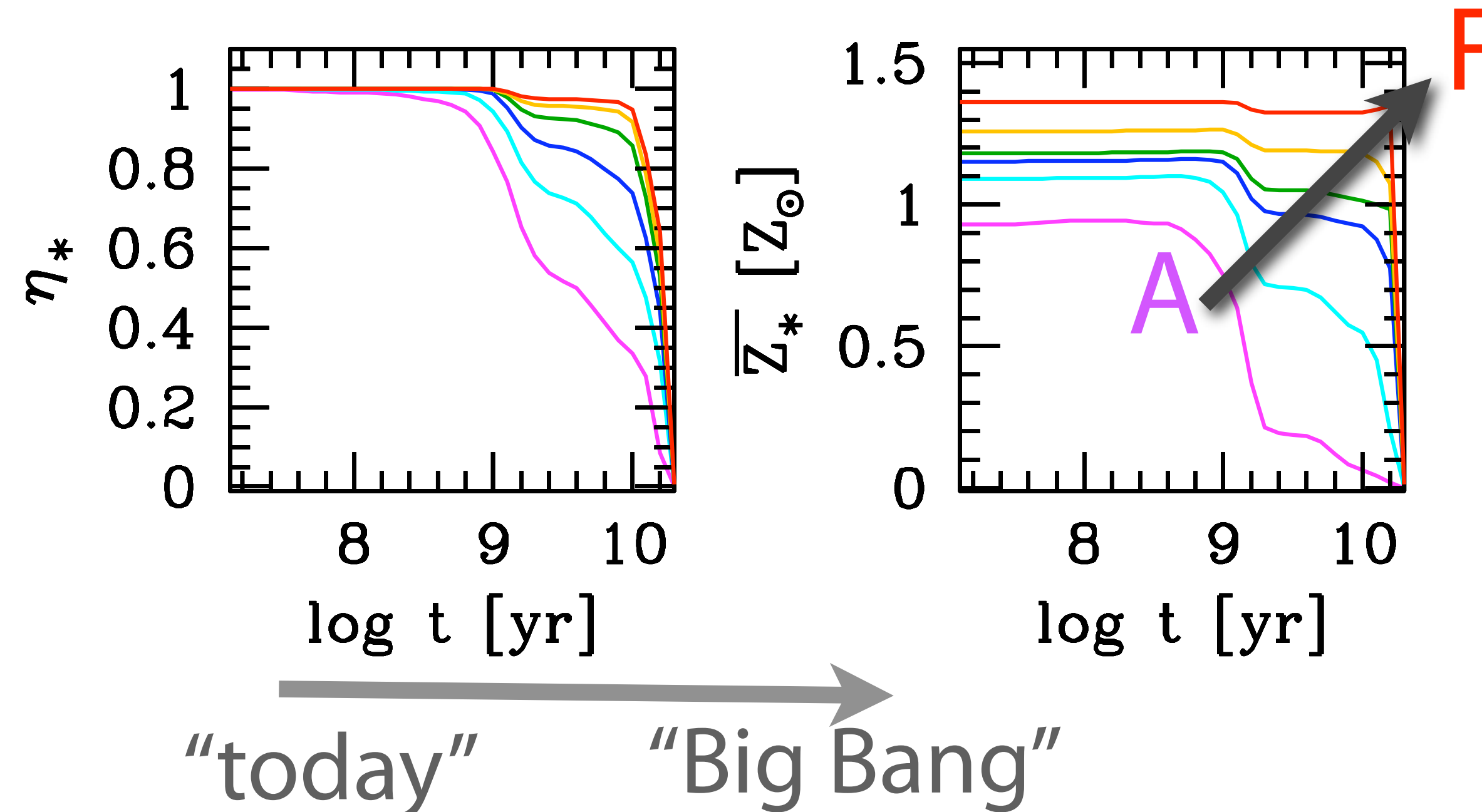


STARLIGHT fits a galaxy spectrum with a sum of simple stellar populations (SSPs) of different ages and metallicities. This is a fit for one galaxy out of the 573141 Sloan Digital Sky Survey objects we have analyzed.

Histories of mass and stellar metallicity

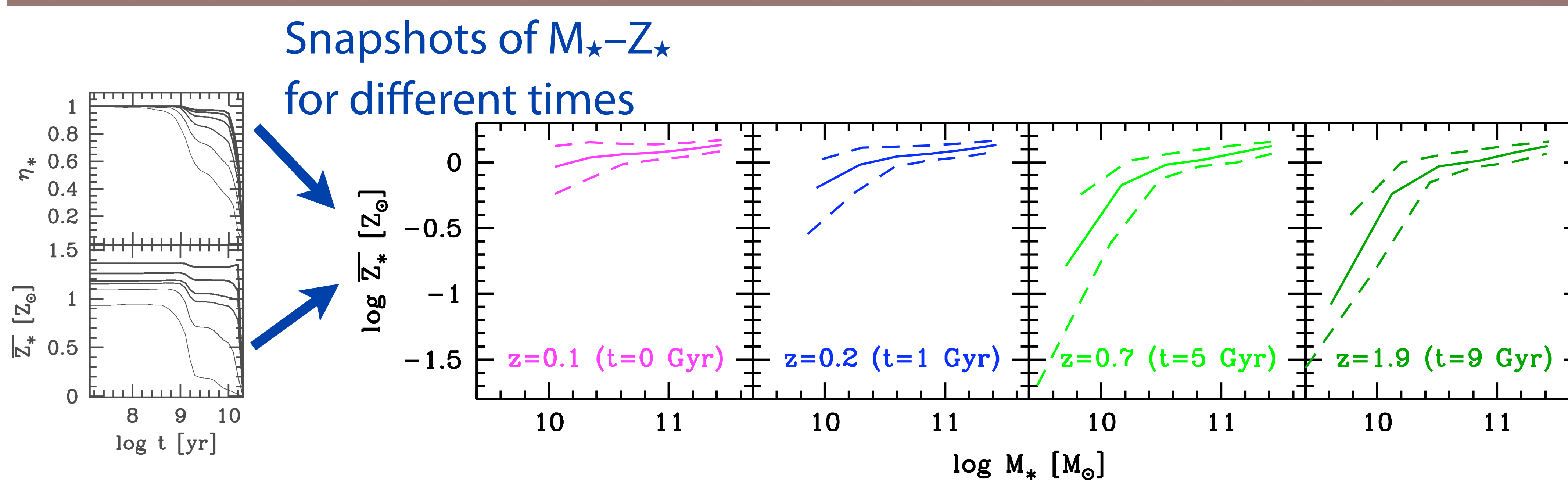
Bins in $\log M/M_{\text{sun}}$, 0.3 dex-wide, centered in:

A: 10.0 D: 10.9
B: 10.3 E: 11.2
C: 10.6 F: 11.5



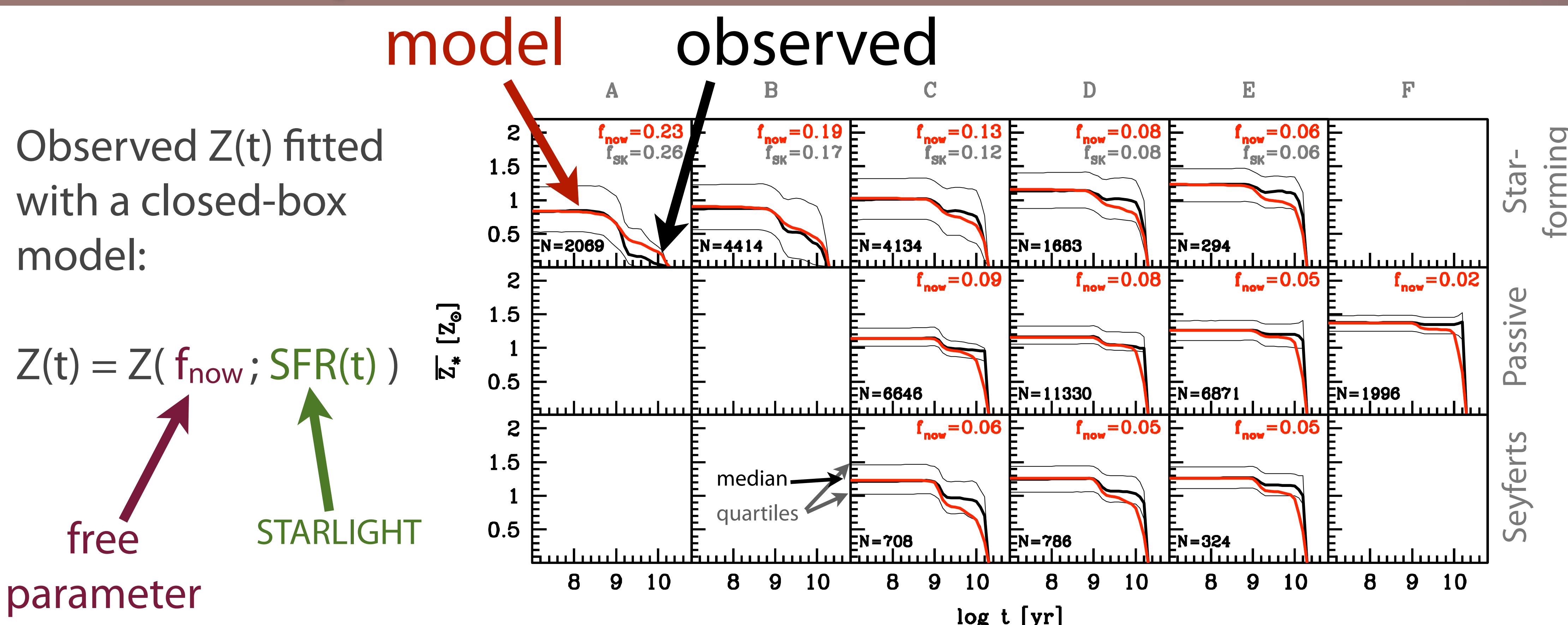
Given the ages and metallicities of the SSPs in a galaxy, we recover the history of the conversion of gas into stars and of metal formation. We group galaxies with similar present-day stellar mass. We find that the *more massive* a galaxy is today, the *faster* it has formed stars and produced metals.

The evolution of the $M_{\star}$ – $Z_{\star}$ relation



Another way to study the evolution of galaxies is to look at a snapshot of the mass vs. metallicity ($M_{\star}$ – $Z_{\star}$) relation for a given look-back time. The snapshots we show are for the *same set* of galaxies and for the *stellar* metallicity. This is the first time such a study is made.

Simple chemical evolution models



Why is there a $M_{\star}$ – $Z_{\star}$ relation? We show that a closed box fits the metallicity histories observed in the range of masses we studied and for different types of objects: Seyfert, passive and star-forming galaxies.

Our results suggest that the $M_{\star}$ – $Z_{\star}$ relation for galaxies with present-day stellar masses down to $10^{10} M_{\text{sun}}$ is mainly *driven by the star formation history* and not by gas inflows or outflows.

More details

* Paper: Vale Asari, N.; Stasinska, G.; Cid Fernandes, R.; Gomes, J. M.; Schlickmann, M.; Mateus, A.; & Schoenell, W. 2009, MNRAS, 396, L71
* STARLIGHT and Virtual Observatory: <http://starlight.ufsc.br/>