

From exoplanets to cosmology
By combining the advantages of angular resolution and/or sensitivity and/or the full spectral range offered by the Moon, one can **For exoplanets** (see TUMP 8)

Carry out spectroscopy of an exoplanet's atmosphere across the full range of wavelengths.

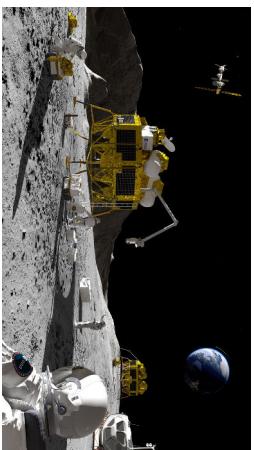
Some exoplanets may have oceans, and if so, we might be able to see the reflected light of the parent star.

For cosmology

Observe the hydrogen line emitted at 21 cm, which for a primordial galaxy with a spectral shift of 30, is observed at 6.3 m, a wavelength a wavelength that is difficult to observe from Earth.

For the first galaxies

Detect the distortion of the primordial radiation spectrum (see TUMP 12) by galaxies on the line of sight.

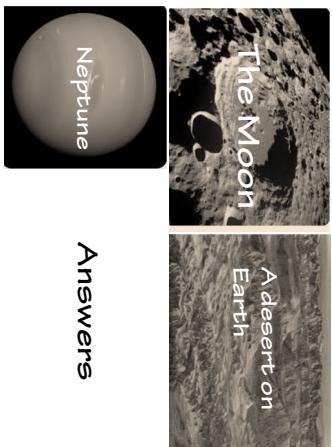


- 1 The Moon offers many practical and astronomical advantages for new observations of the stars.
- 2 Scientific goals and spin-offs
- 3 Implementation



A volcano
on Venus

Gravity on the Moon is 1/6th of that of the Earth



Answers



To learn more about this series and about the topics presented in this booklet, please visit:
<http://www.tump.org>

Translation:StanKurtz
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Why the Moon?

On the Moon there is no atmosphere, so no atmospheric turbulence, and we have access to the entire light spectrum, from gamma rays to radio waves.

The gravity which is 1/6th of Earth's gravity means that one can build there telescopes that are larger than on Earth. In addition the instruments can be constantly upgraded and repaired.

To set up and operate telescopes on the Moon, one needs to know the terrain, send robots and then humans, and build houses. One will also need vehicles to transport people from one place to another as well as energy sources.

Investigations on all these aspects are progressing very quickly.

The Universe in my pocket No 40

This mini-book was written in 2025 by Jean Schneider of the Paris Observatory and revised by Stan Kurtz of the IRyA (Mexico).

It gives an instant view of a subject that is progressing very rapidly: see luth7.obspm.fr/Moon.html

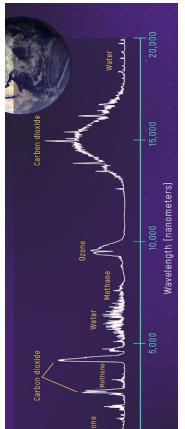
The cover image is a representation of future astronomical activity on the Moon, with its optical and radio telescopes and the astronomers and engineers operating the instruments.

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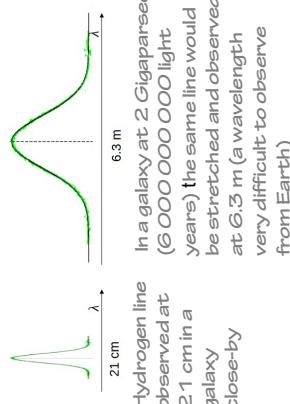
Completed and planned telescopes

- **Apoll0 16 (1972)**
- **A 7.5 cm telescope**
- **Lunar Ultraviolet Telescope (2013)**
- **A 15 cm telescope**
- **ROSES (2024)**
- **Radio telescope**
- **LOUPE**
- **A 30 cm telescope**
- **LISTER dedicated to the spectroscopy of exoplanet transits**
- **Grat telescope**

- For transit spectroscopy and imaging
- As sophisticated camera to observe the Earth as if it were an exoplanet
- For the detection of all stars colder than 100 degrees Celsius
- Long-base interferometers
- General radio telescopes
- For the detection of long wavelength stars



We can see the entire spectrum (see TUMP 30) of an exoplanet, from the ultra-violet to the far infrared.



In a galaxy at 2 Gigaparsec (600,000,000 light years), the same line would be observed at 6.3 m (a wavelength close-by 21 cm in a galaxy observed at 21 cm from Earth) very difficult to observe from Earth.